

**Evidence for Extrusive Mg-Suite Magmatism on the Moon? Fine-Grained  
Magnesian Clasts in an Apollo 16 Impact Melt Breccia**

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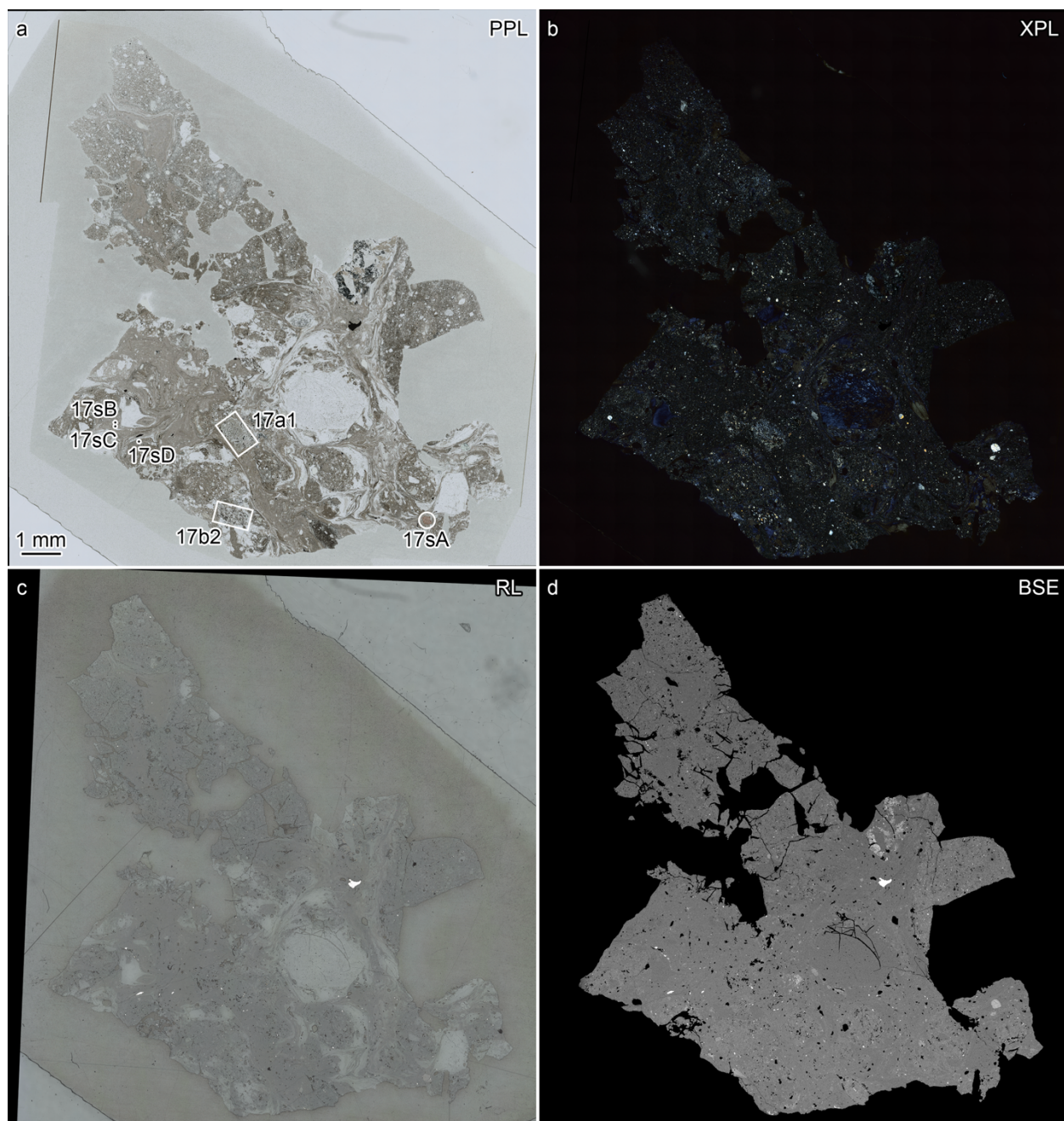


Figure S1. (a) Plane-polarized light (PPL) image of 68815,17 with labeled locations of clasts of interest. (b) Cross-polarized light (XPL) image of 68815,17. (c) Reflected light (RL) image of 68815,17. (d) Backscattered electron (BSE) image of 68815,17.

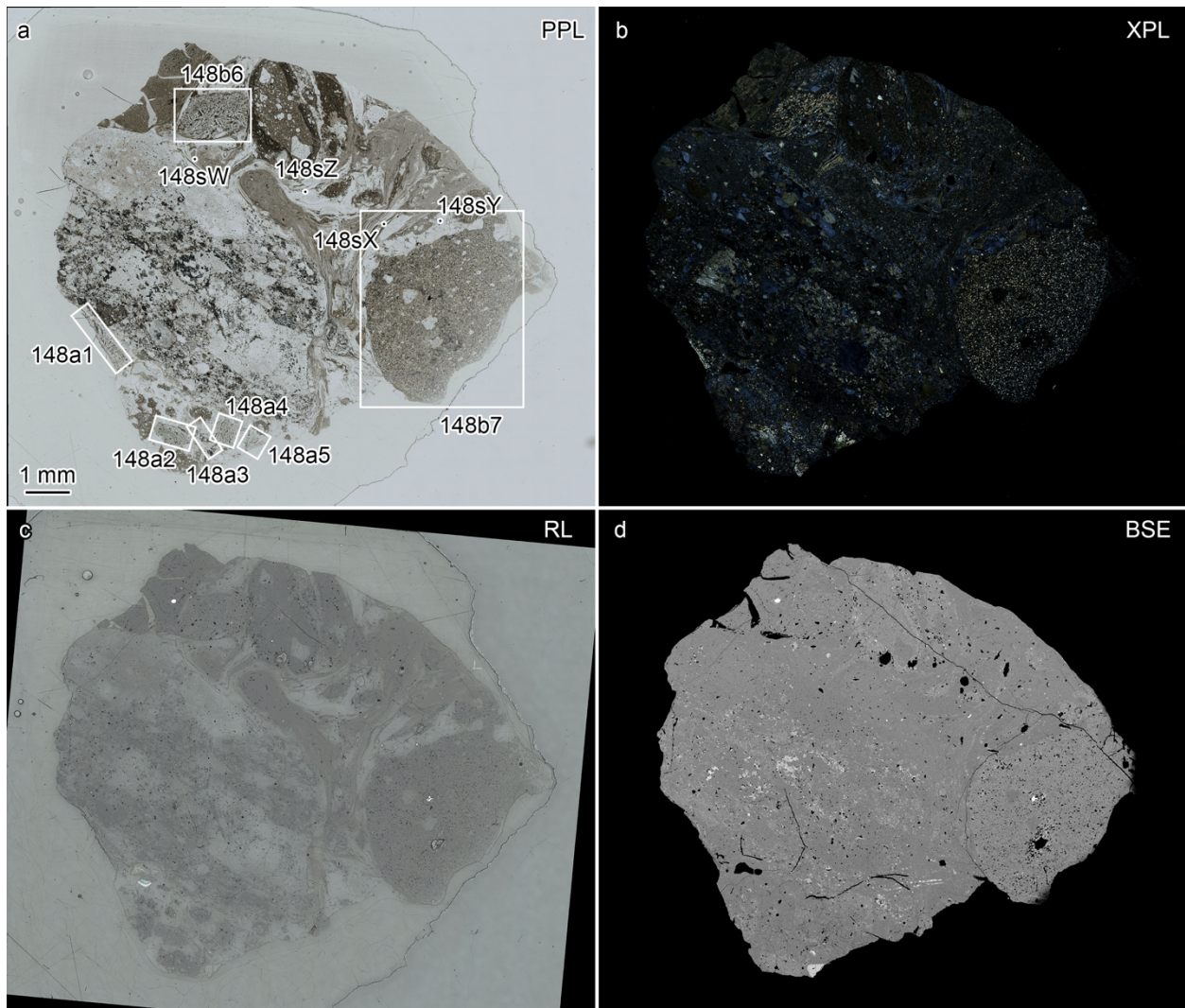


Figure S2. (a) PPL image of 68815,148 with labeled locations of clasts of interest. (b) XPL image of 68815,148. (c) RL image of 68815,148. (d) BSE image of 68815,148.

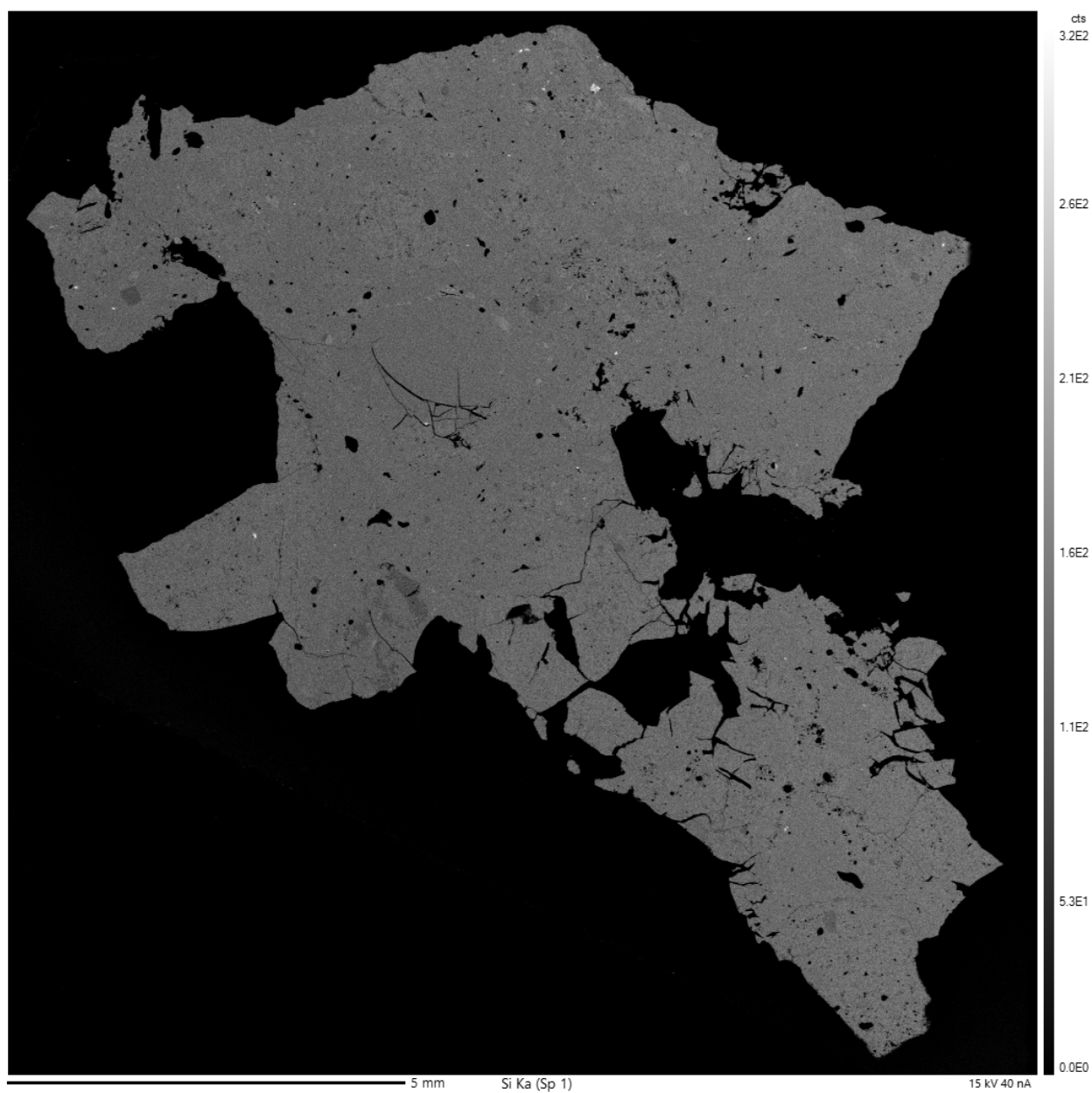


Figure S3. Si X-ray map of 68815,17.

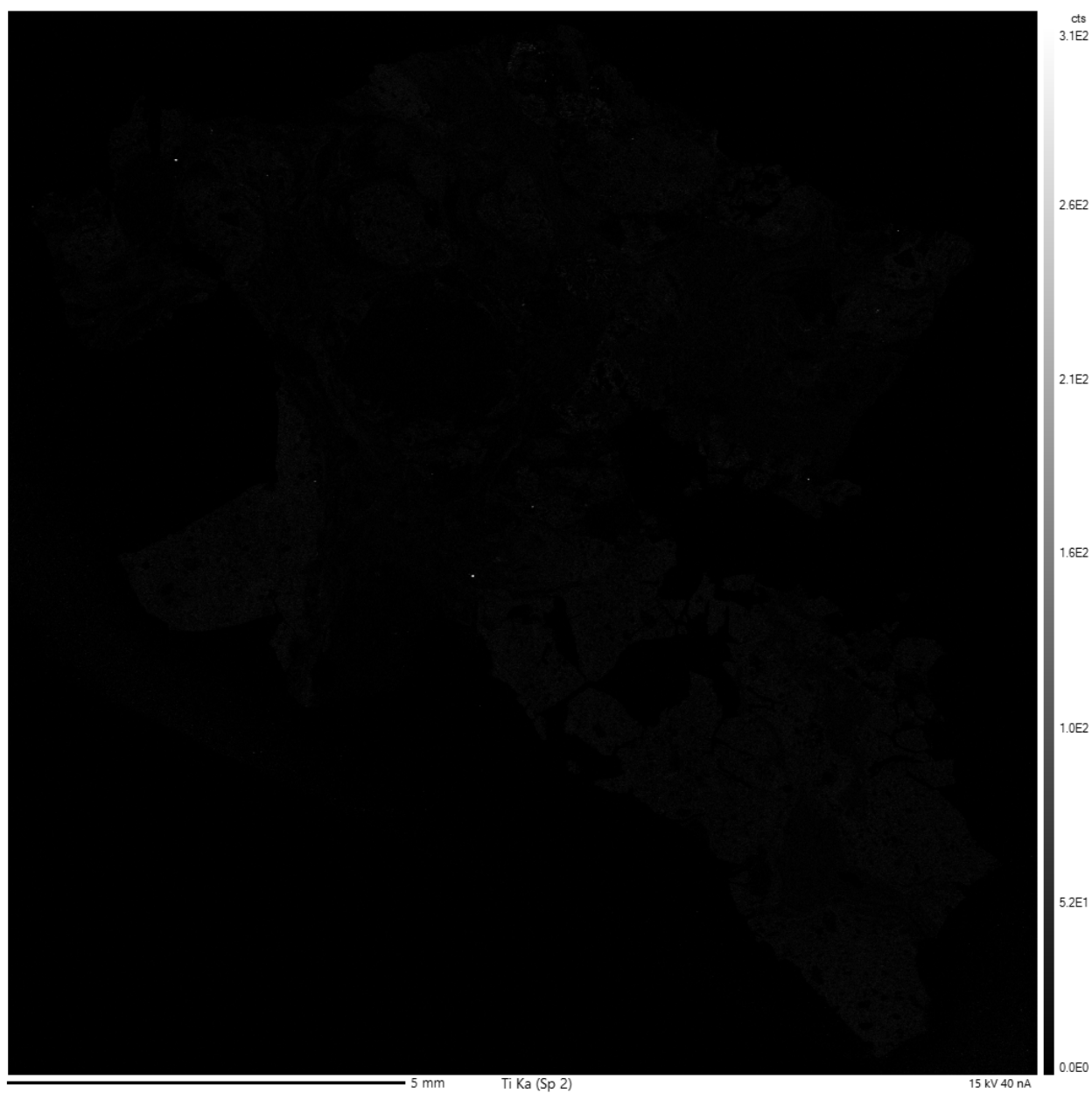


Figure S4. Ti X-ray map of 68815,17.

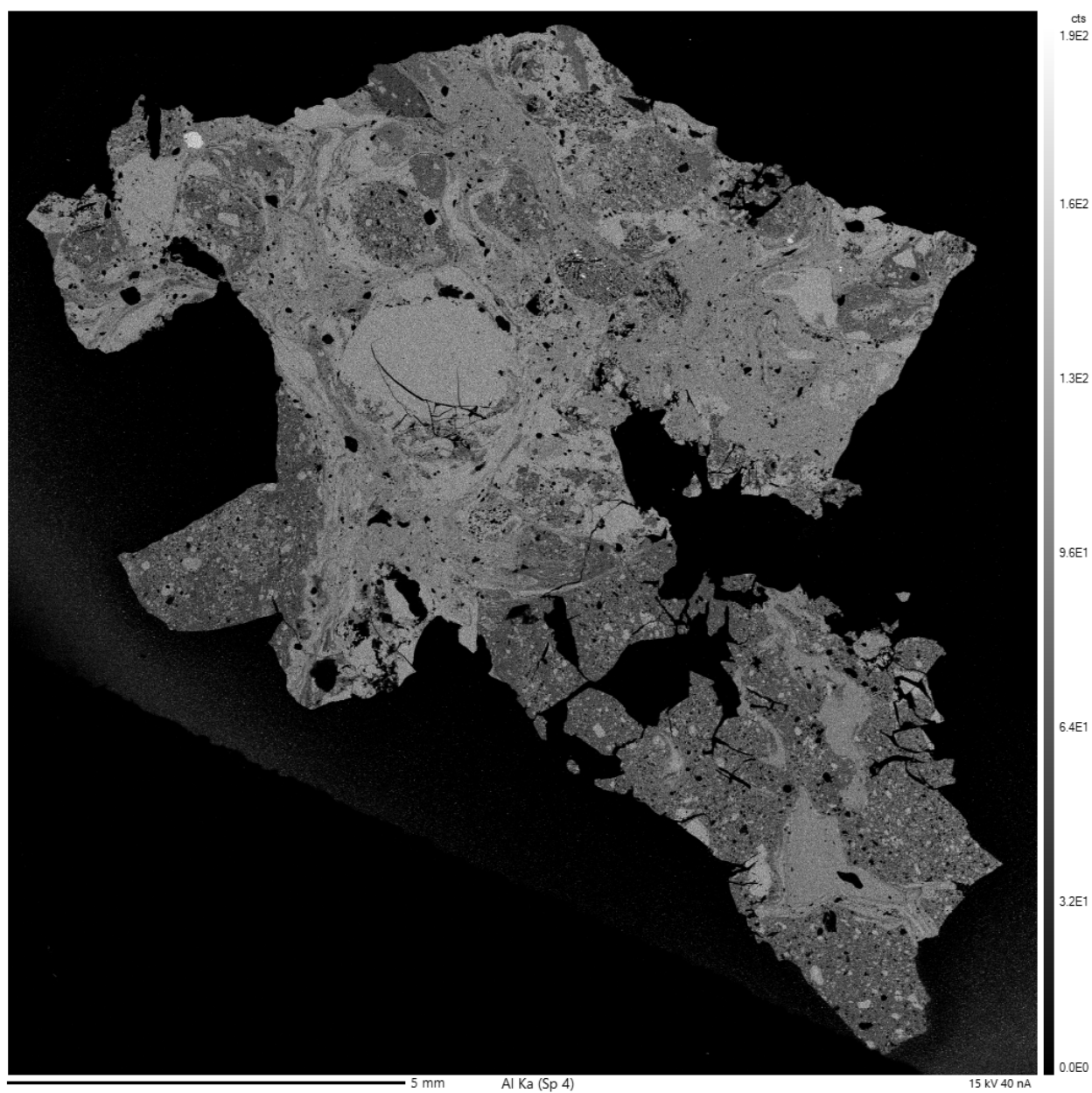


Figure S5. Al X-ray map of 68815,17.

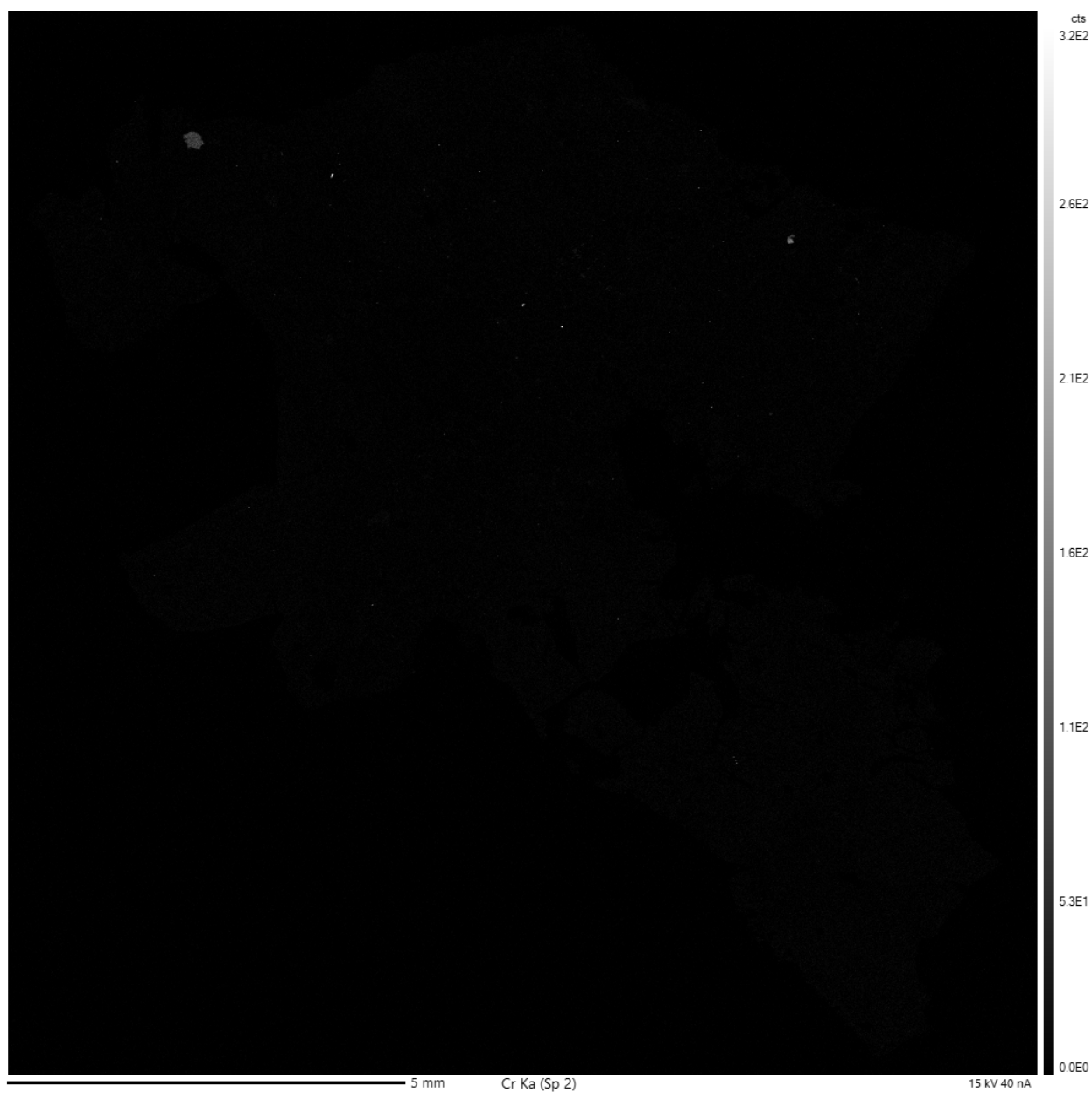


Figure S6. Cr X-ray map of 68815,17.

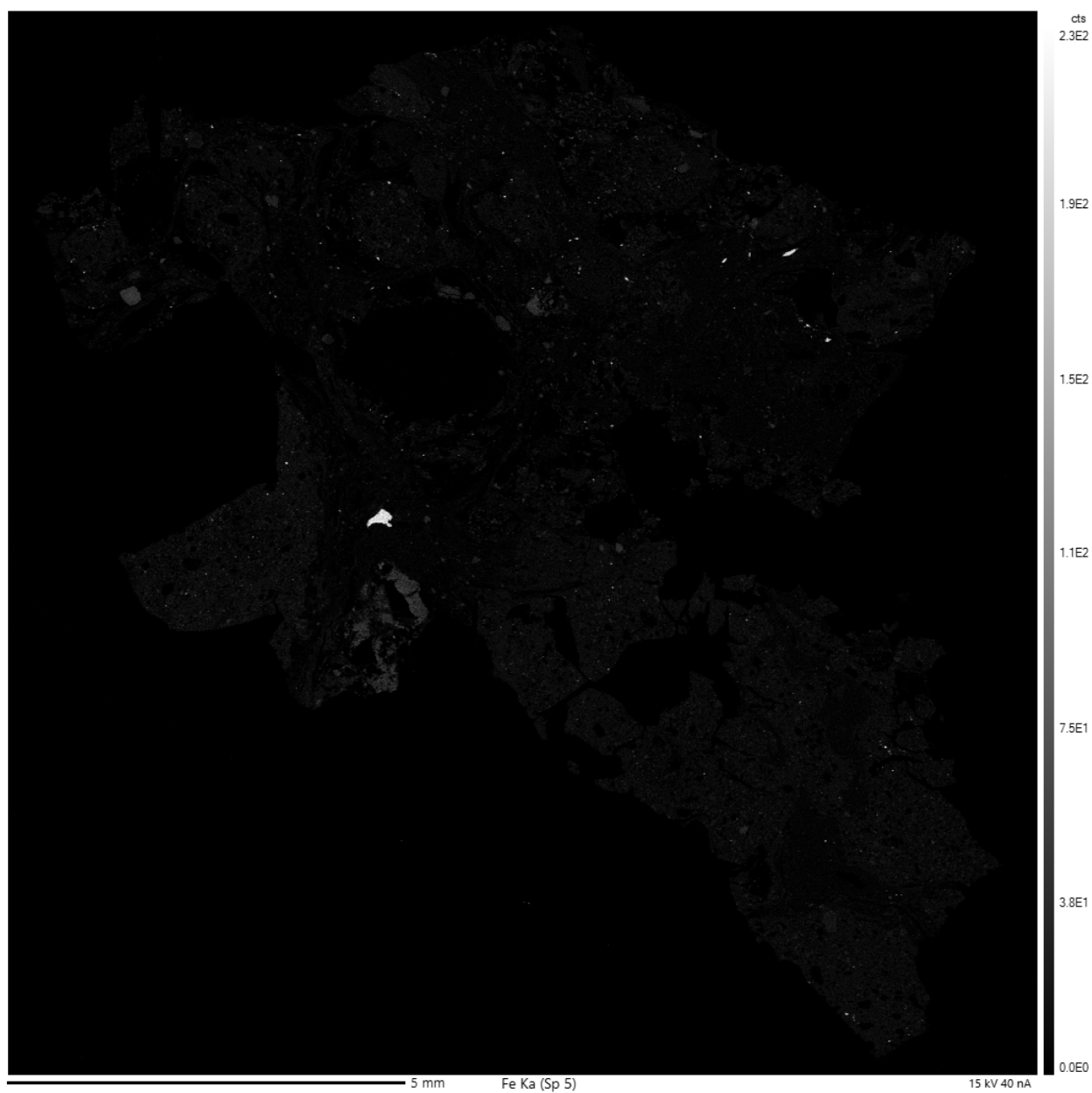


Figure S7. Fe X-ray map of 68815,17.

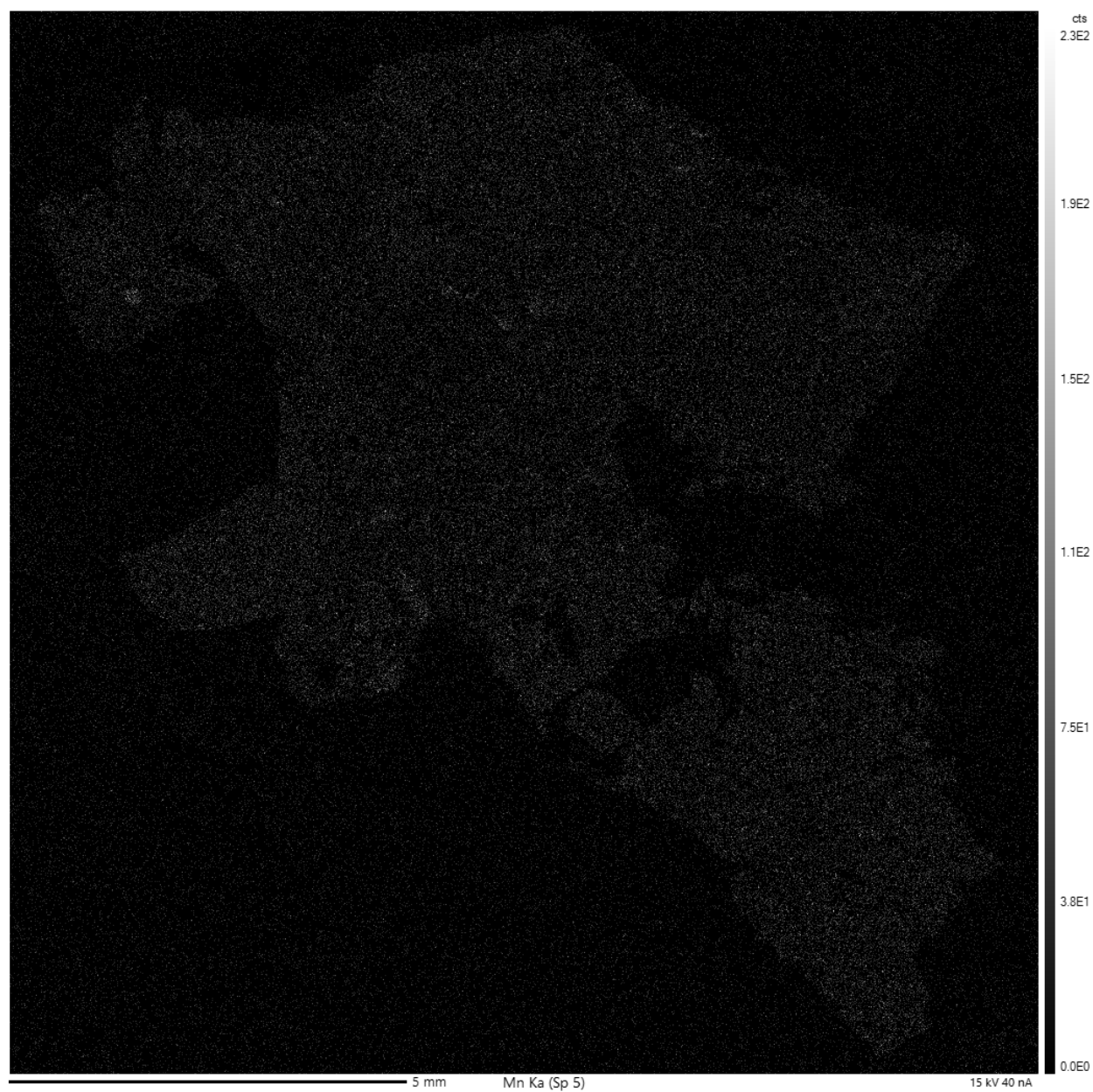


Figure S8. Mn X-ray map of 68815,17.

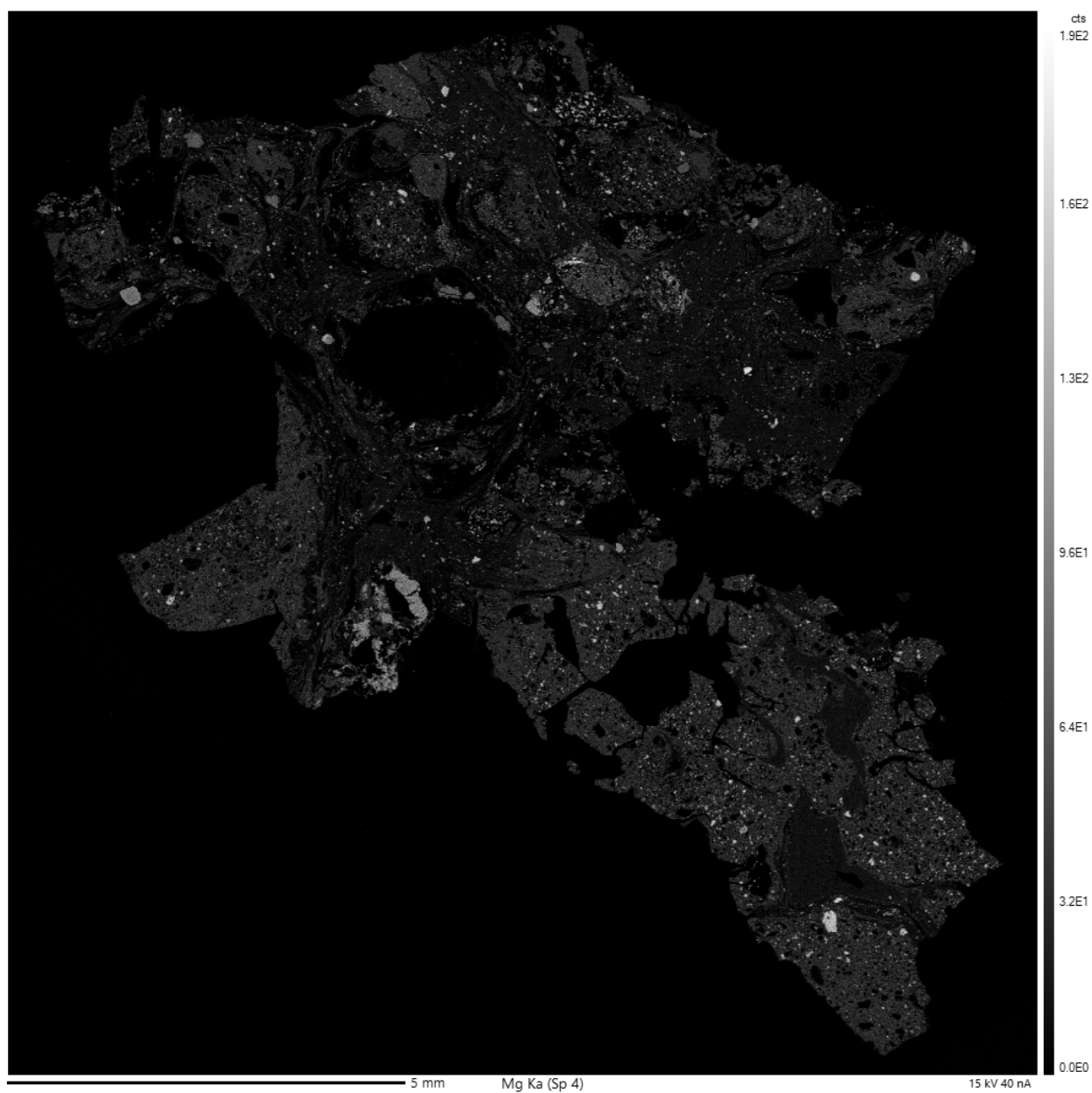


Figure S9. Mg X-ray map of 68815,17.

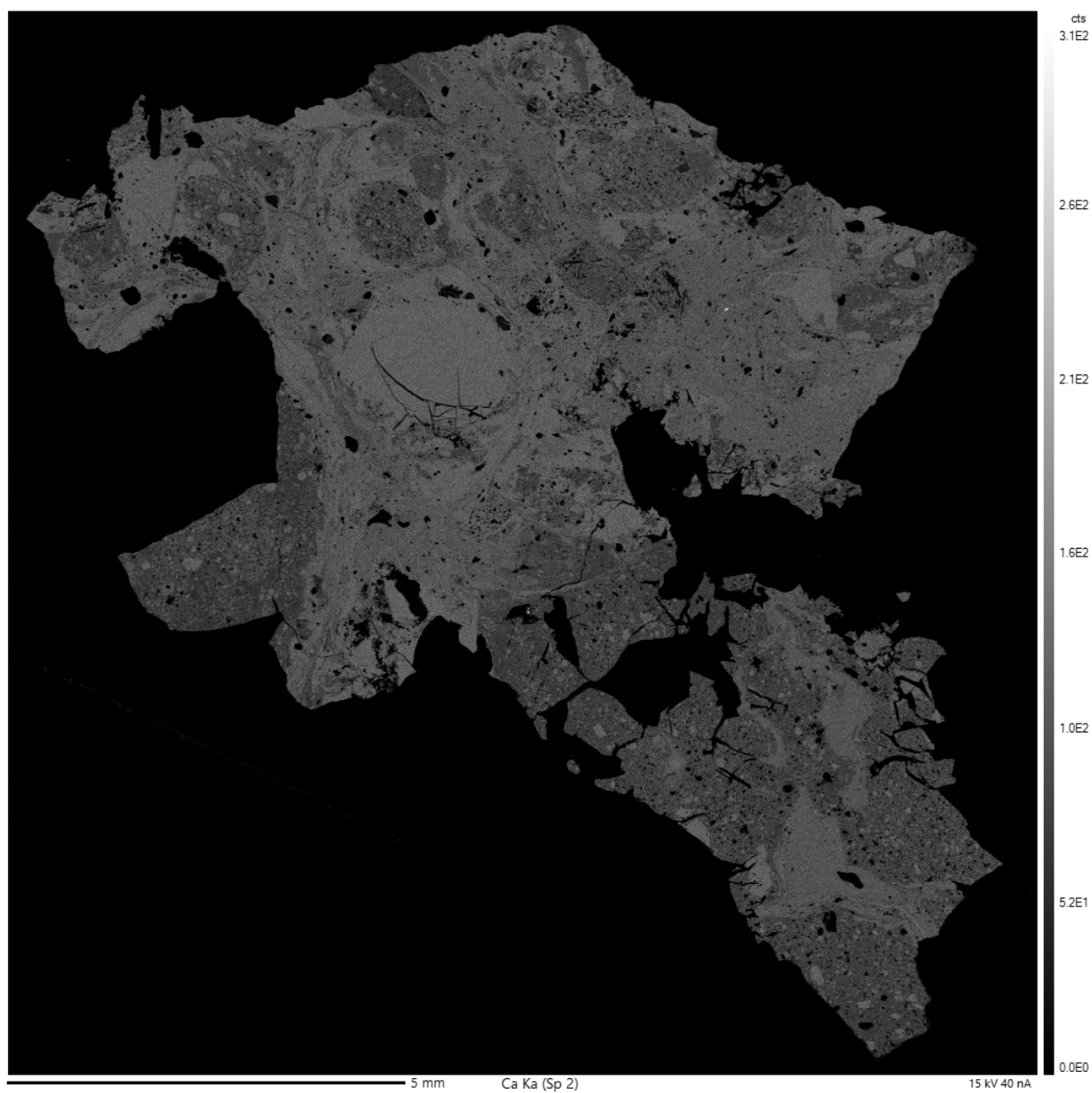


Figure S10. Ca X-ray map of 68815,17.

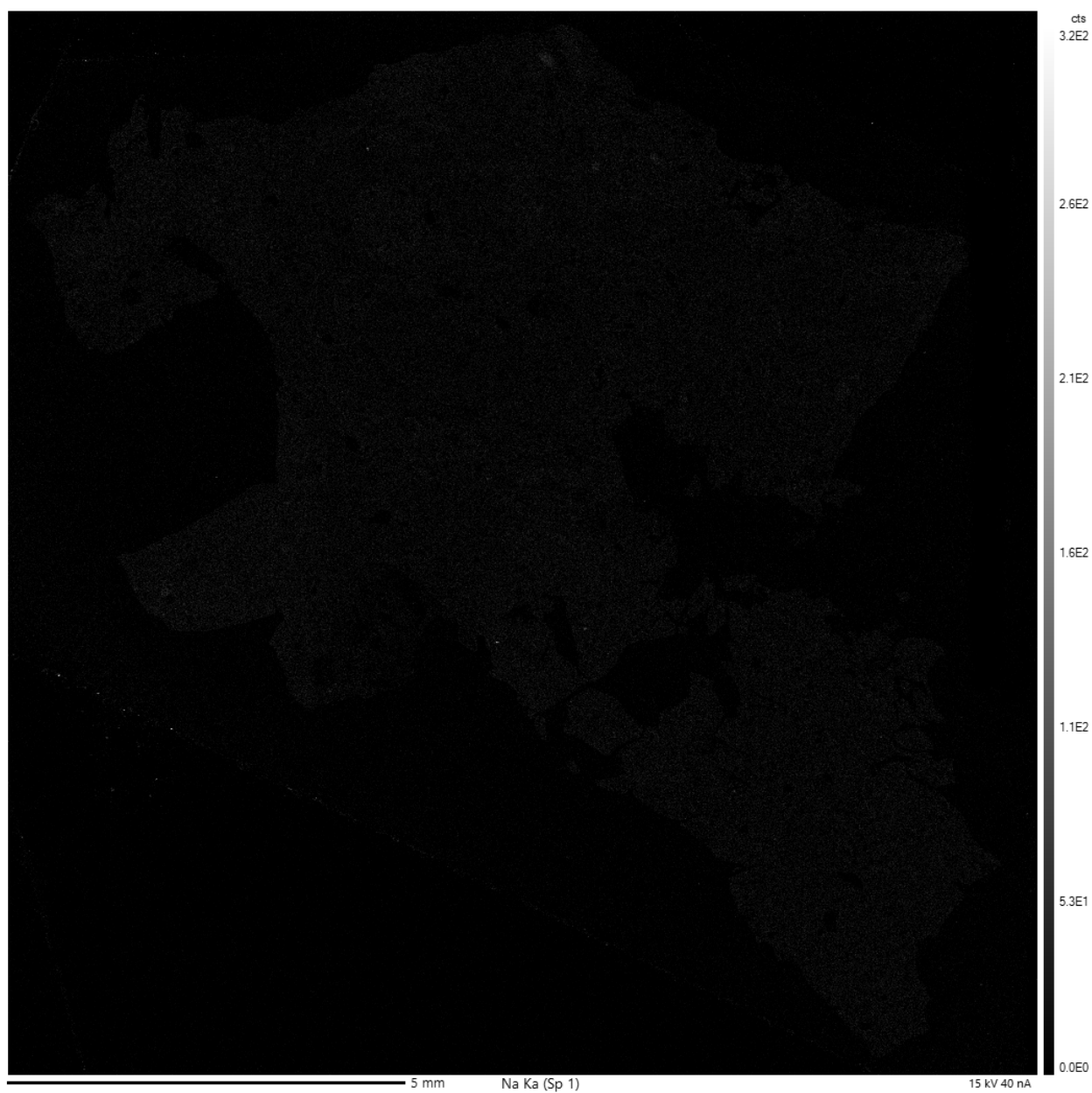


Figure S11. Na X-ray map of 68815,17.

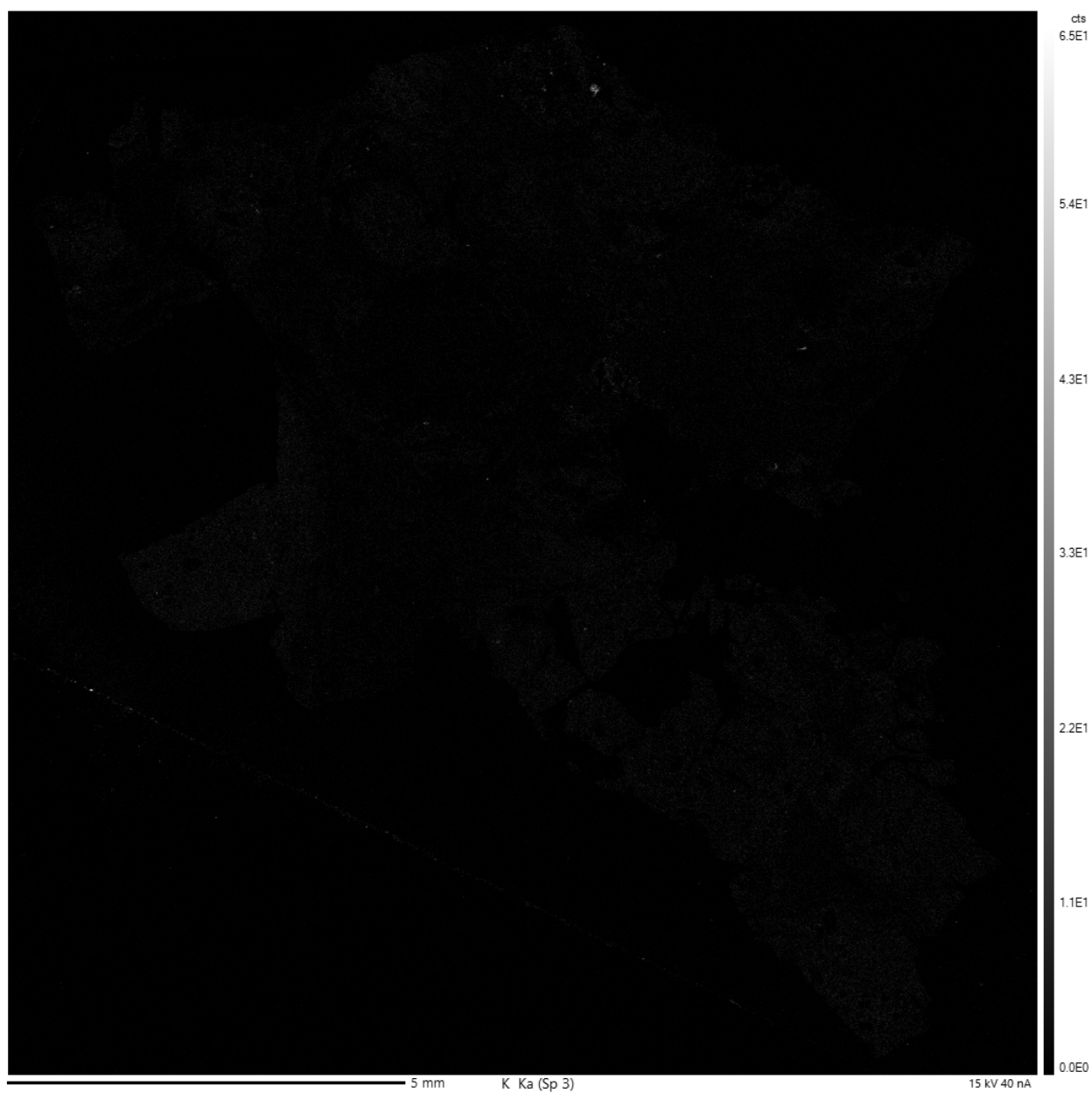


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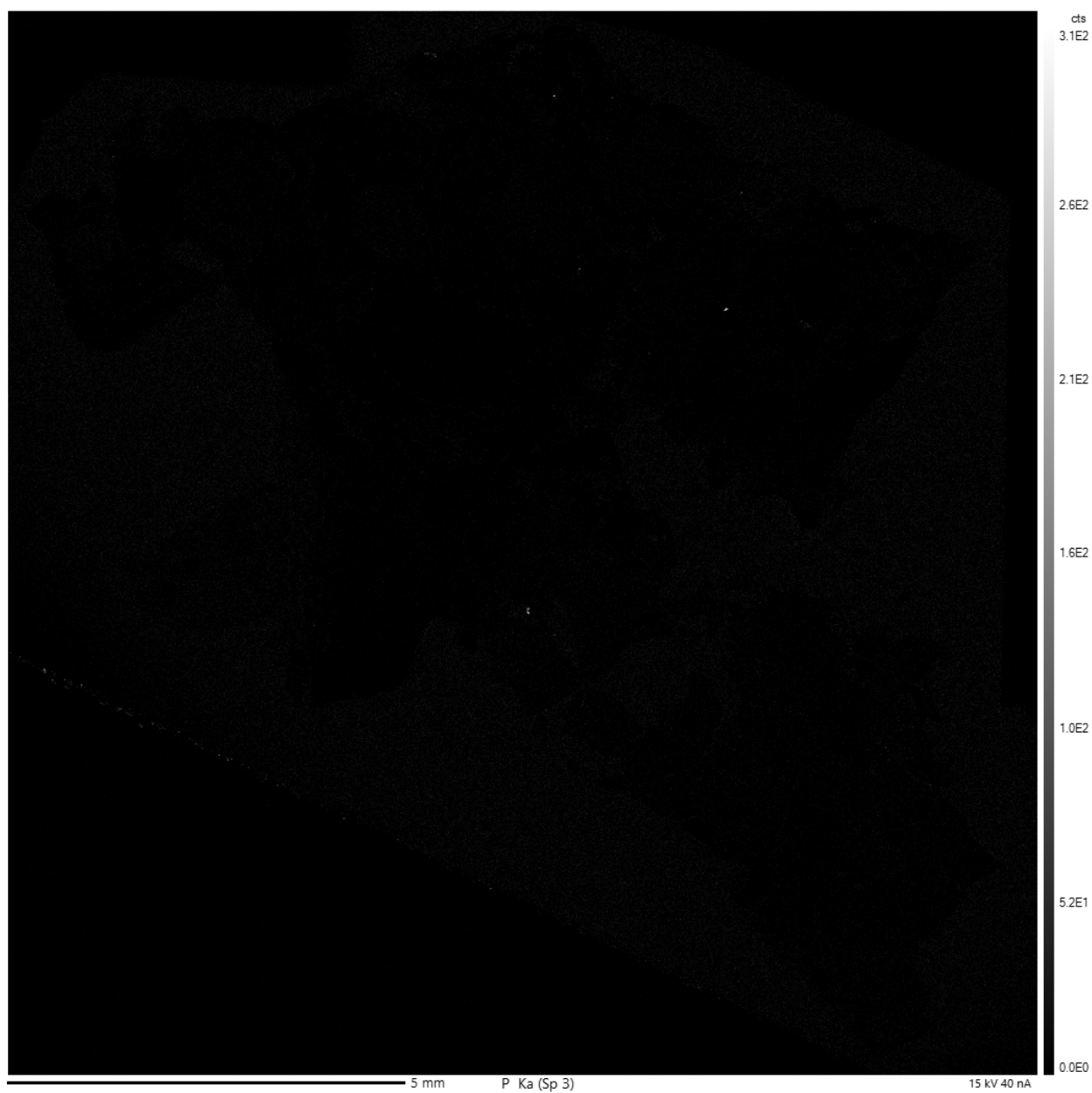


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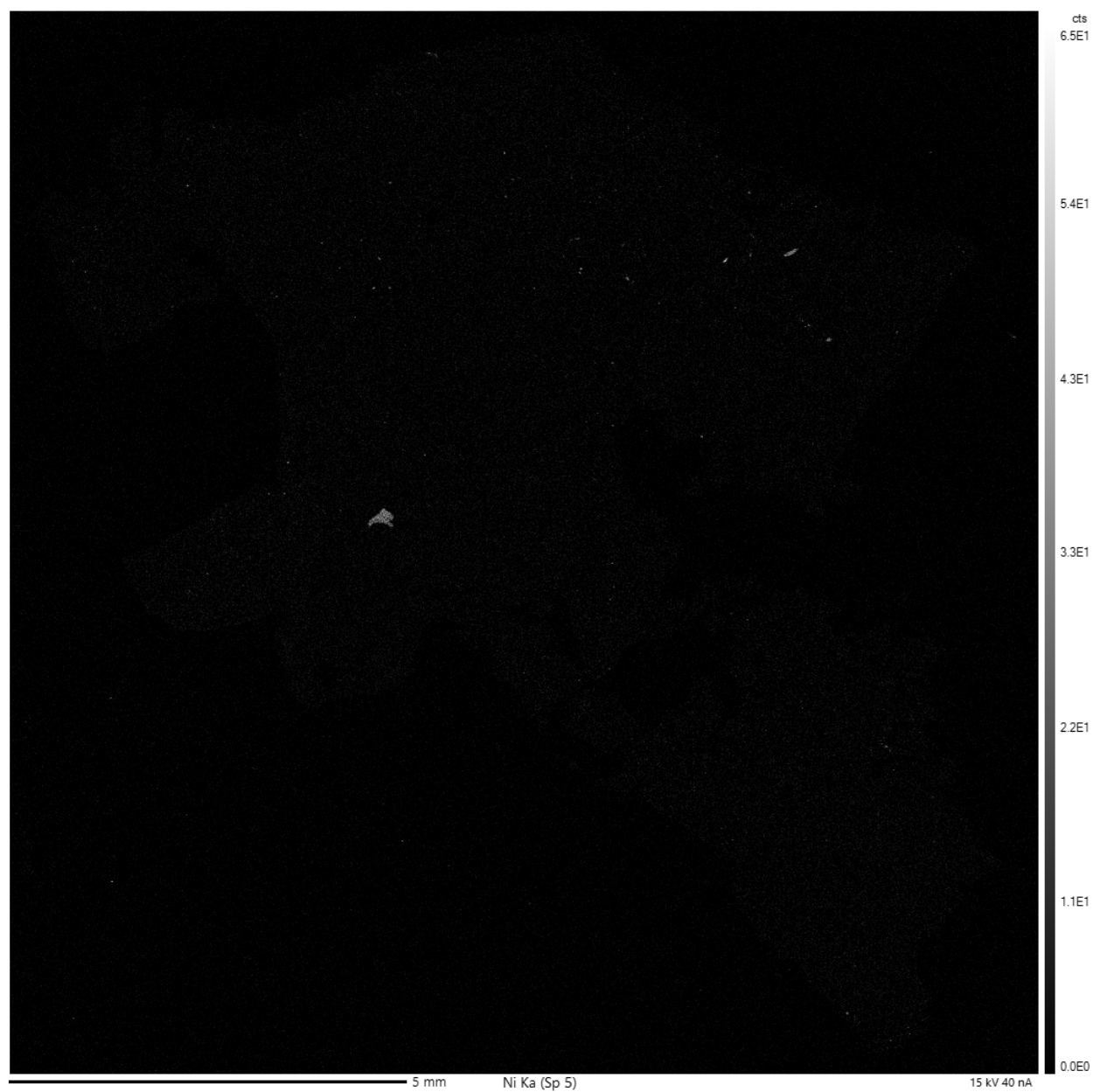


Figure S14. Ni X-ray map of 68815,17.

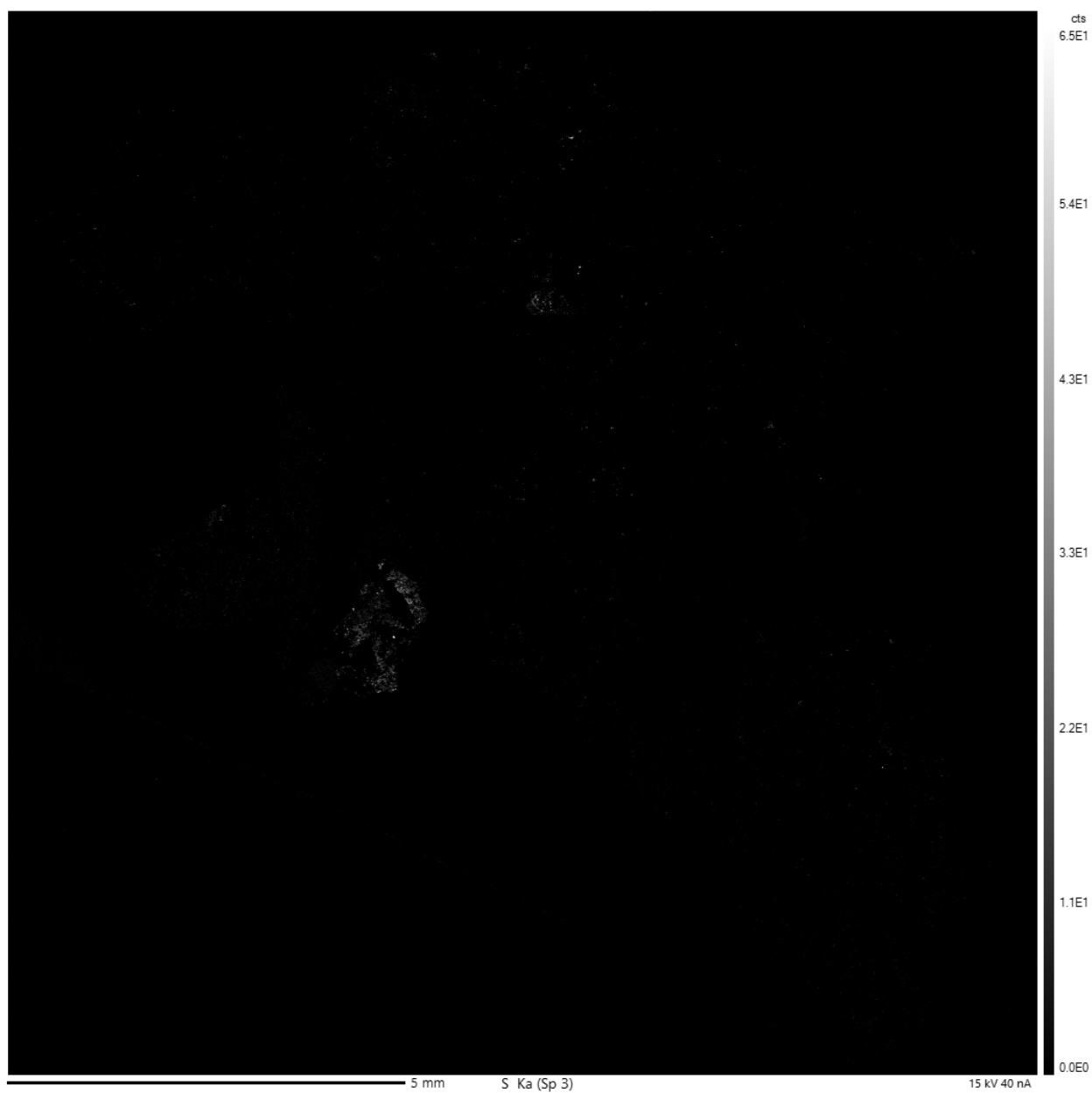


Figure S15. S X-ray map of 68815,17.

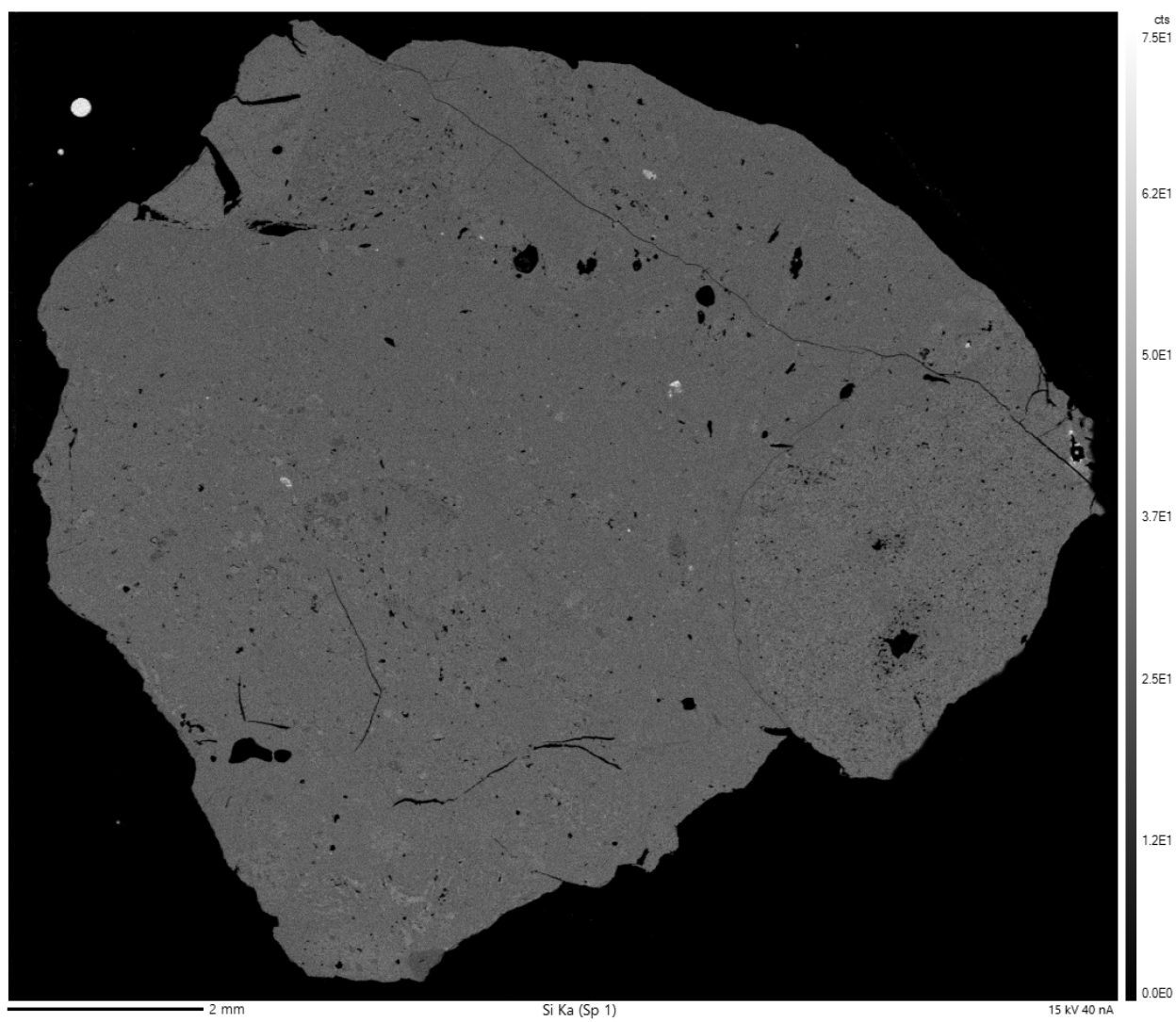


Figure S16. Si X-ray map of 68815,148.

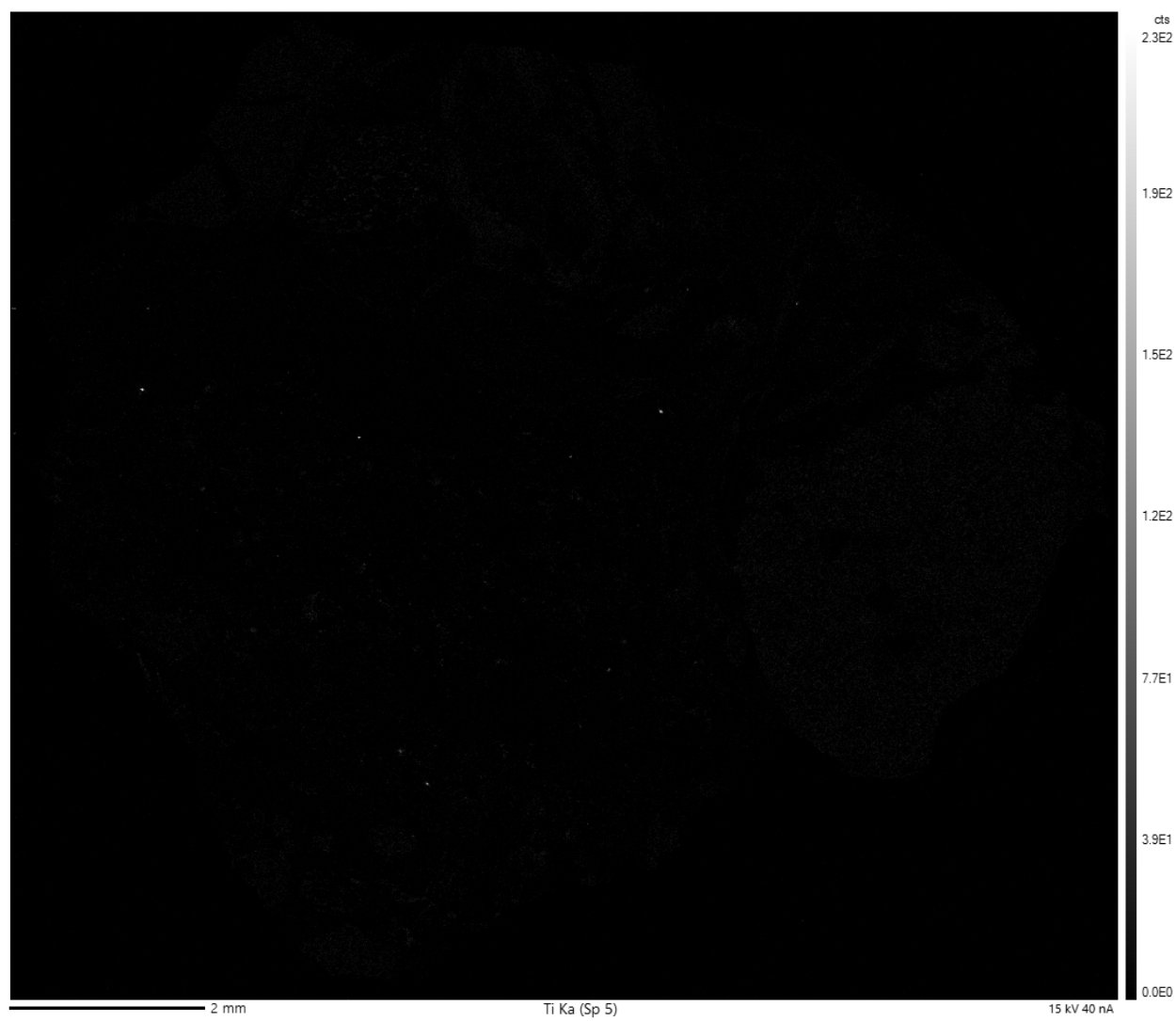


Figure S17. Ti X-ray map of 68815,148.

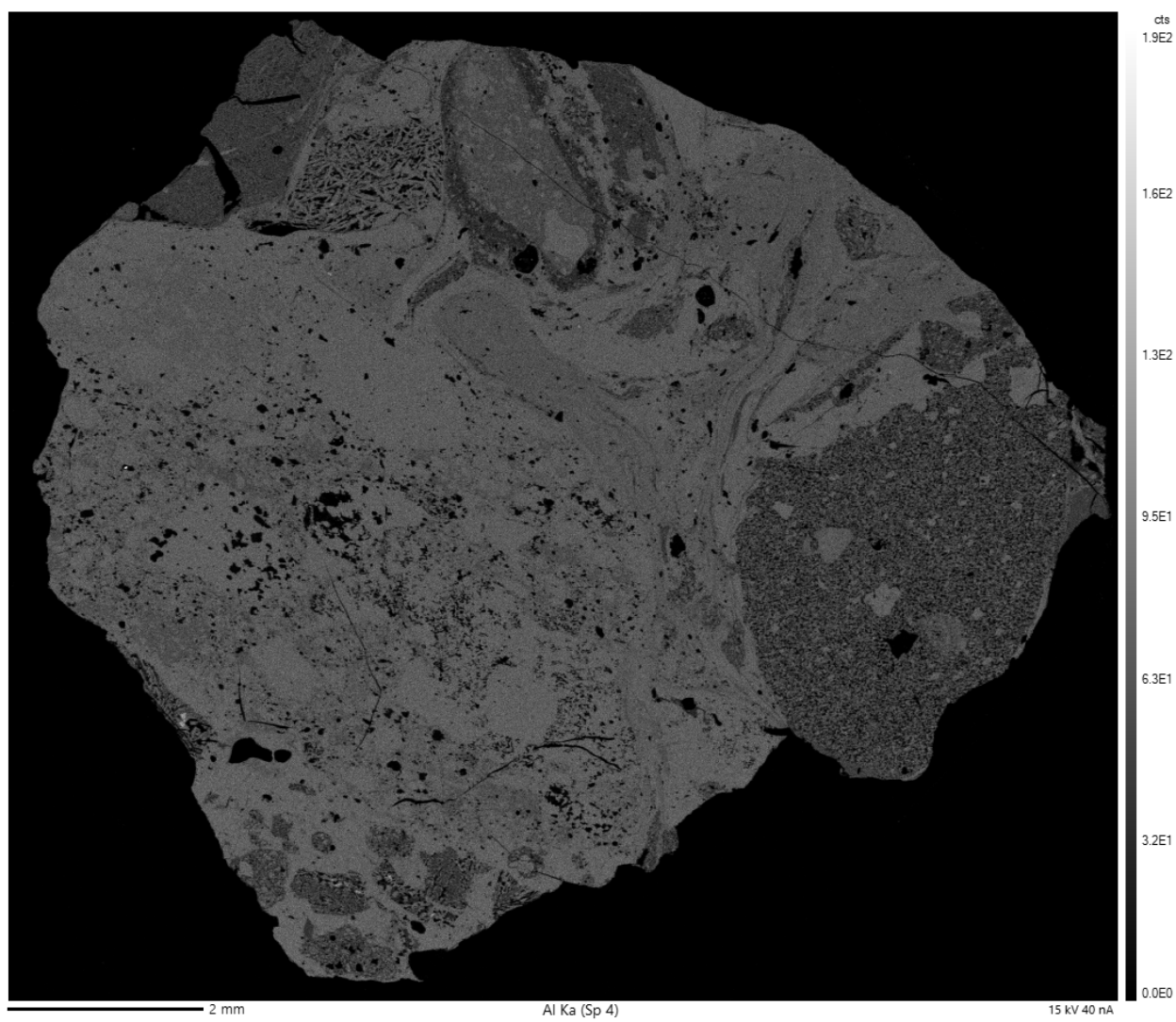


Figure S18. Al X-ray map of 68815,148.



Figure S19. Cr X-ray map of 68815,148.

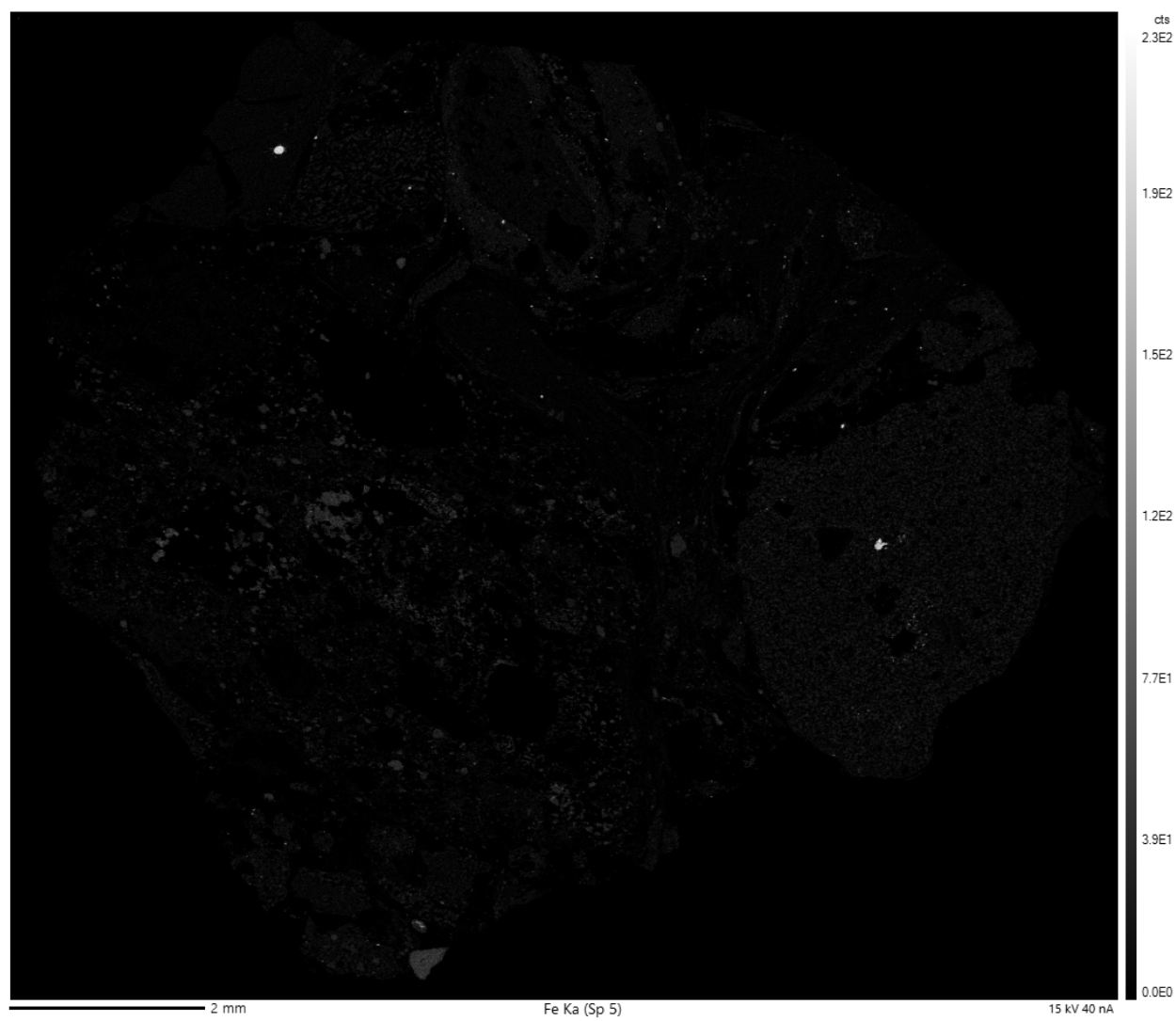


Figure S20. Fe X-ray map of 68815,148.

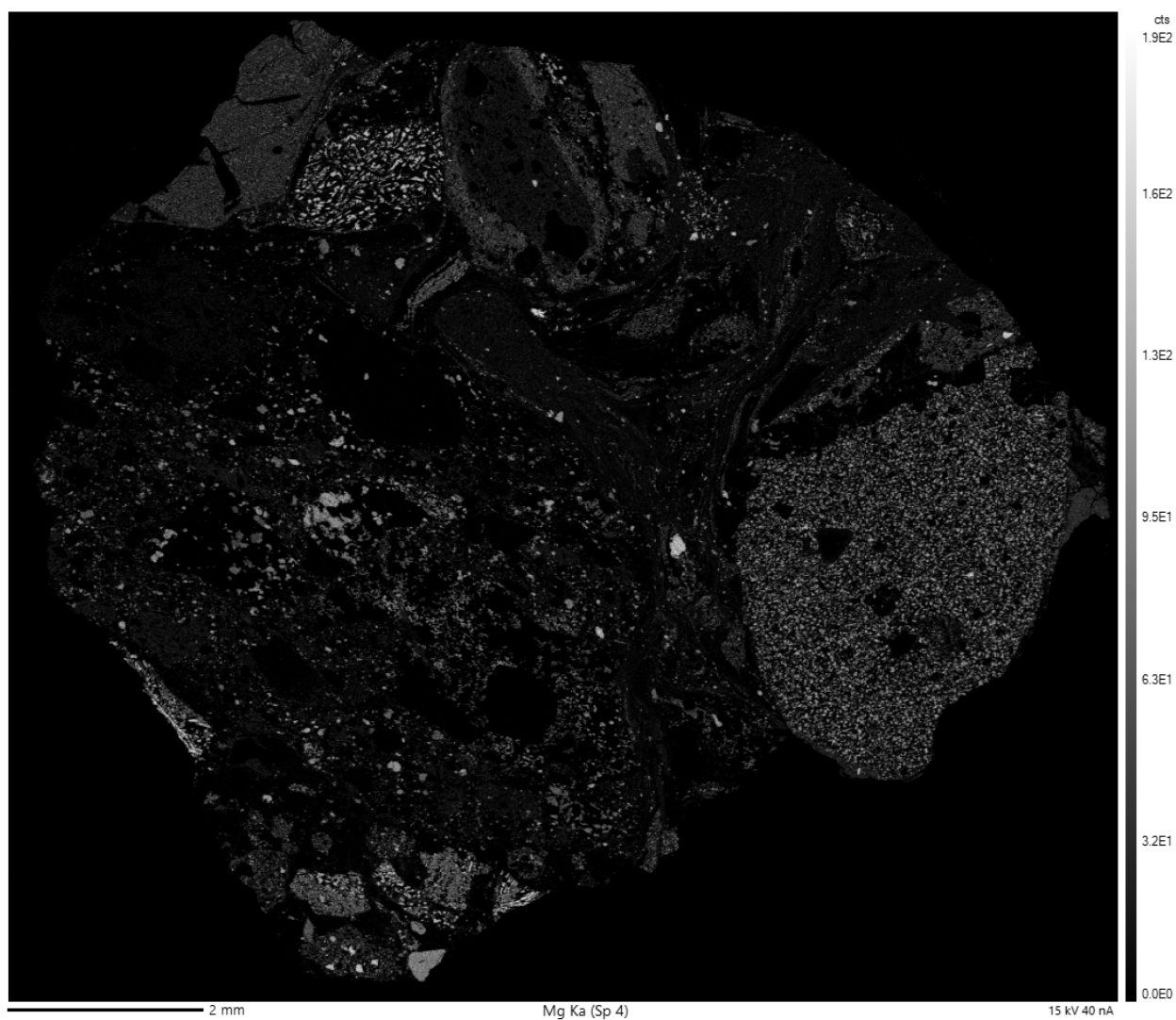


Figure S21. Mg X-ray map of 68815,148.

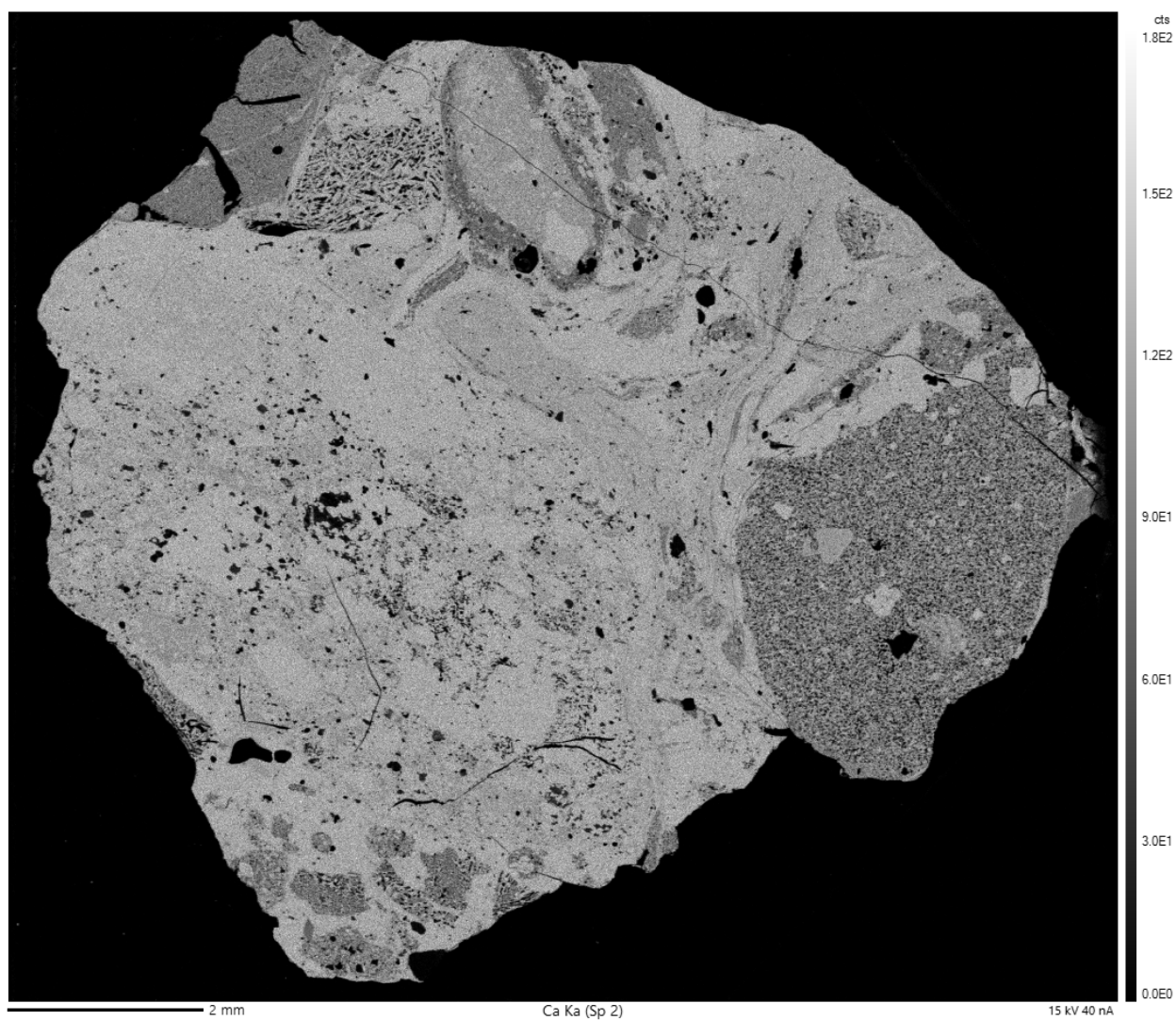


Figure S22. Ca X-ray map of 68815,148.

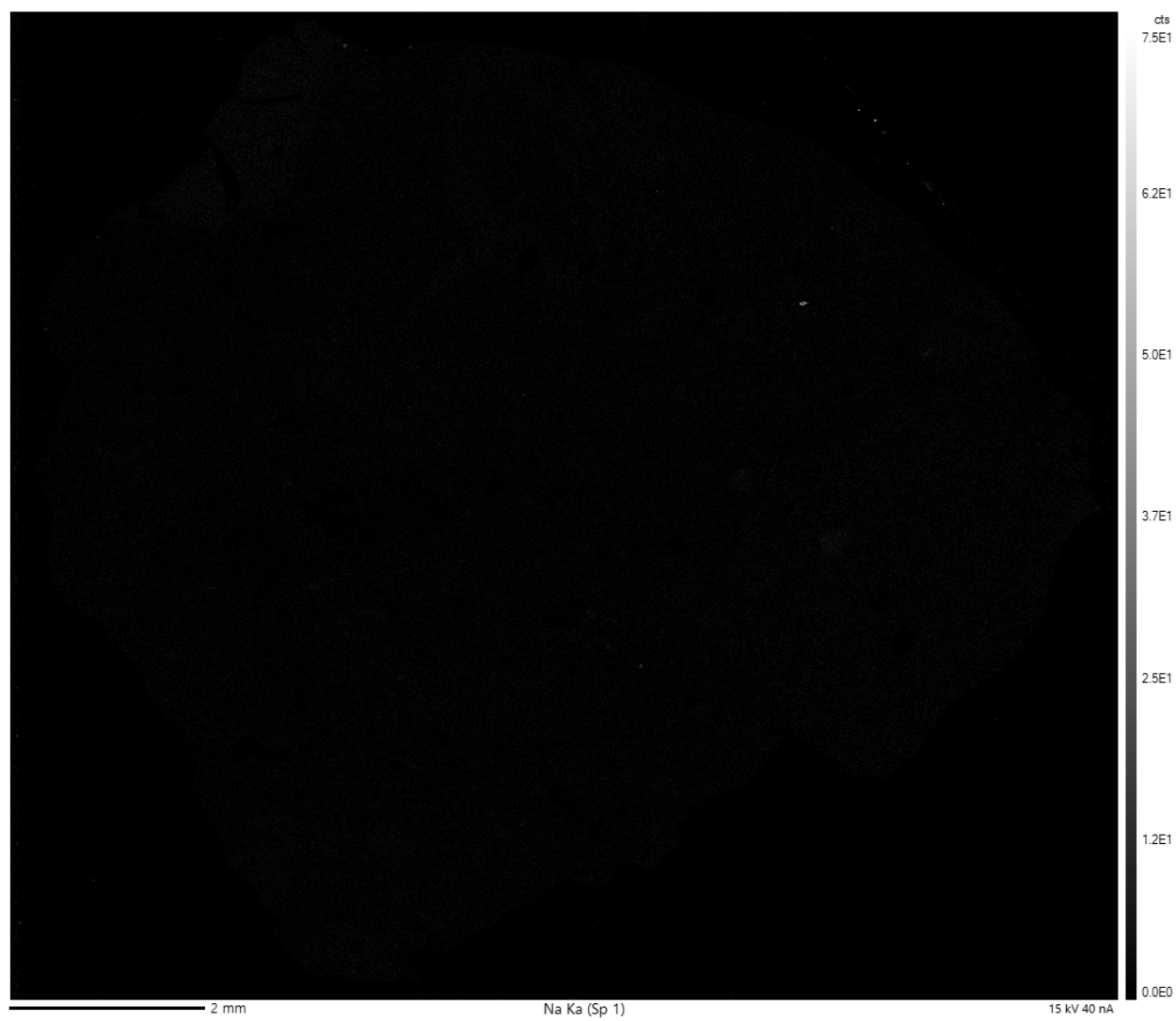


Figure S23. Na X-ray map of 68815,148.

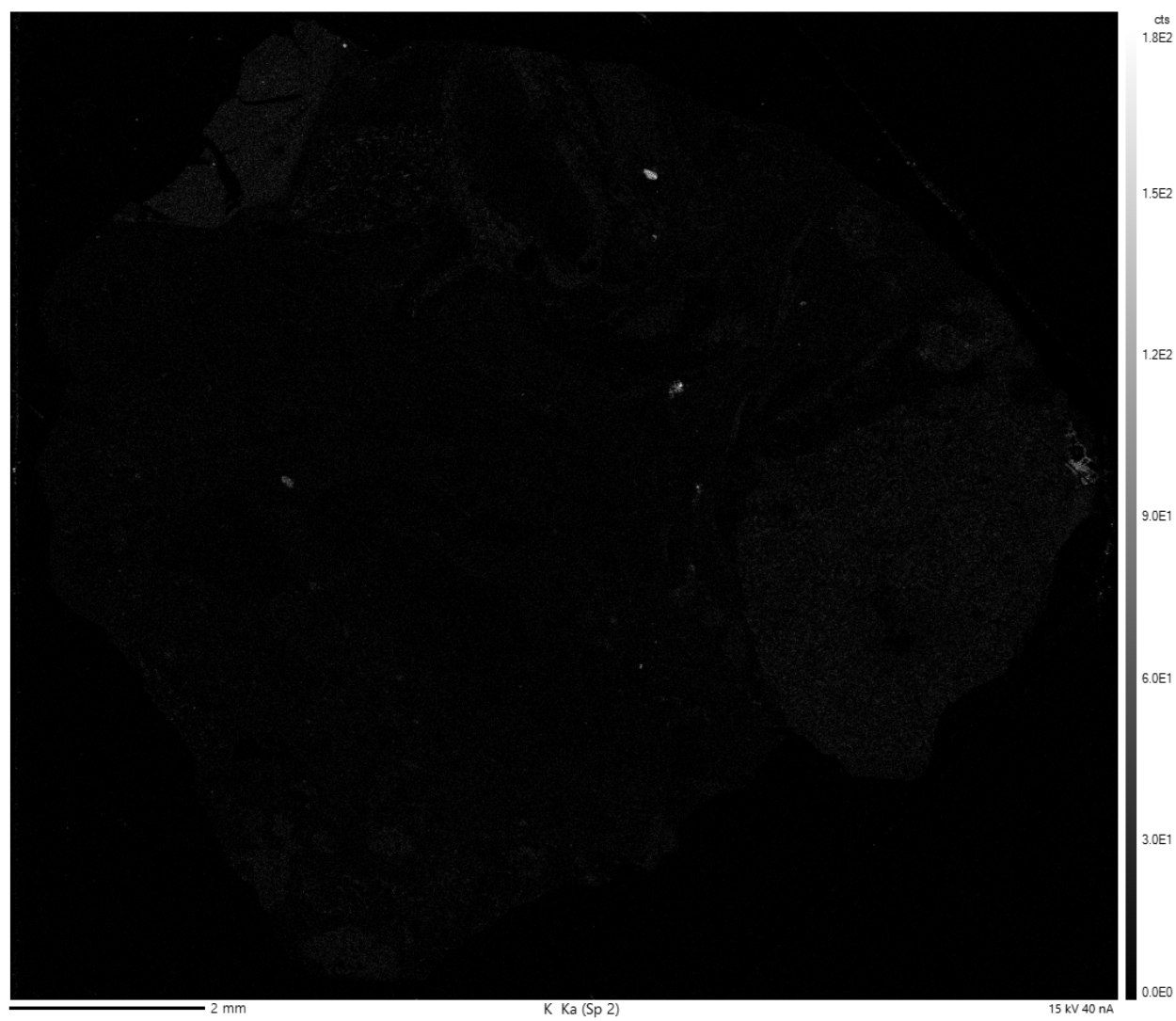


Figure S24. K X-ray map of 68815,148.

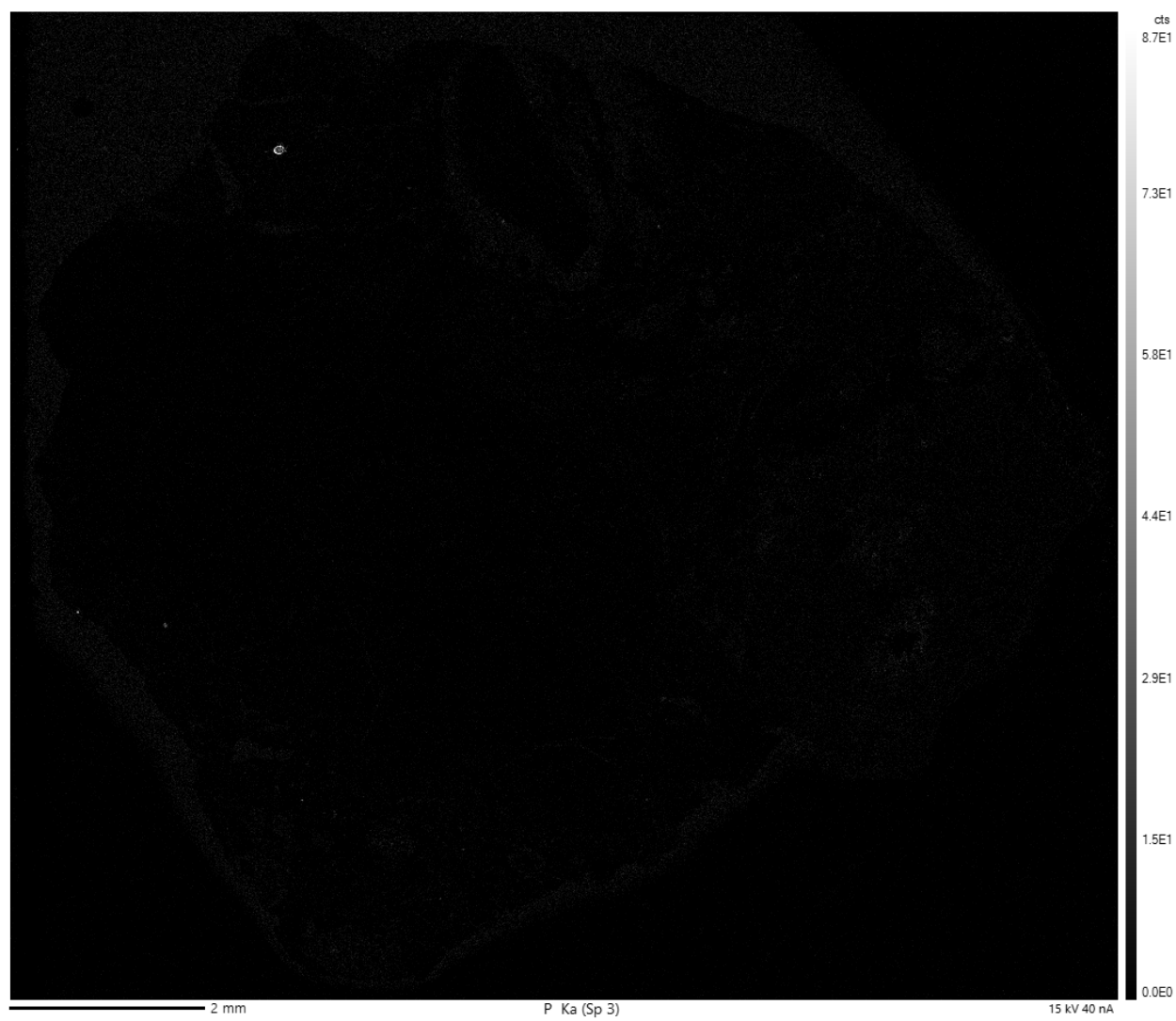


Figure S25. P X-ray map of 68815,148.

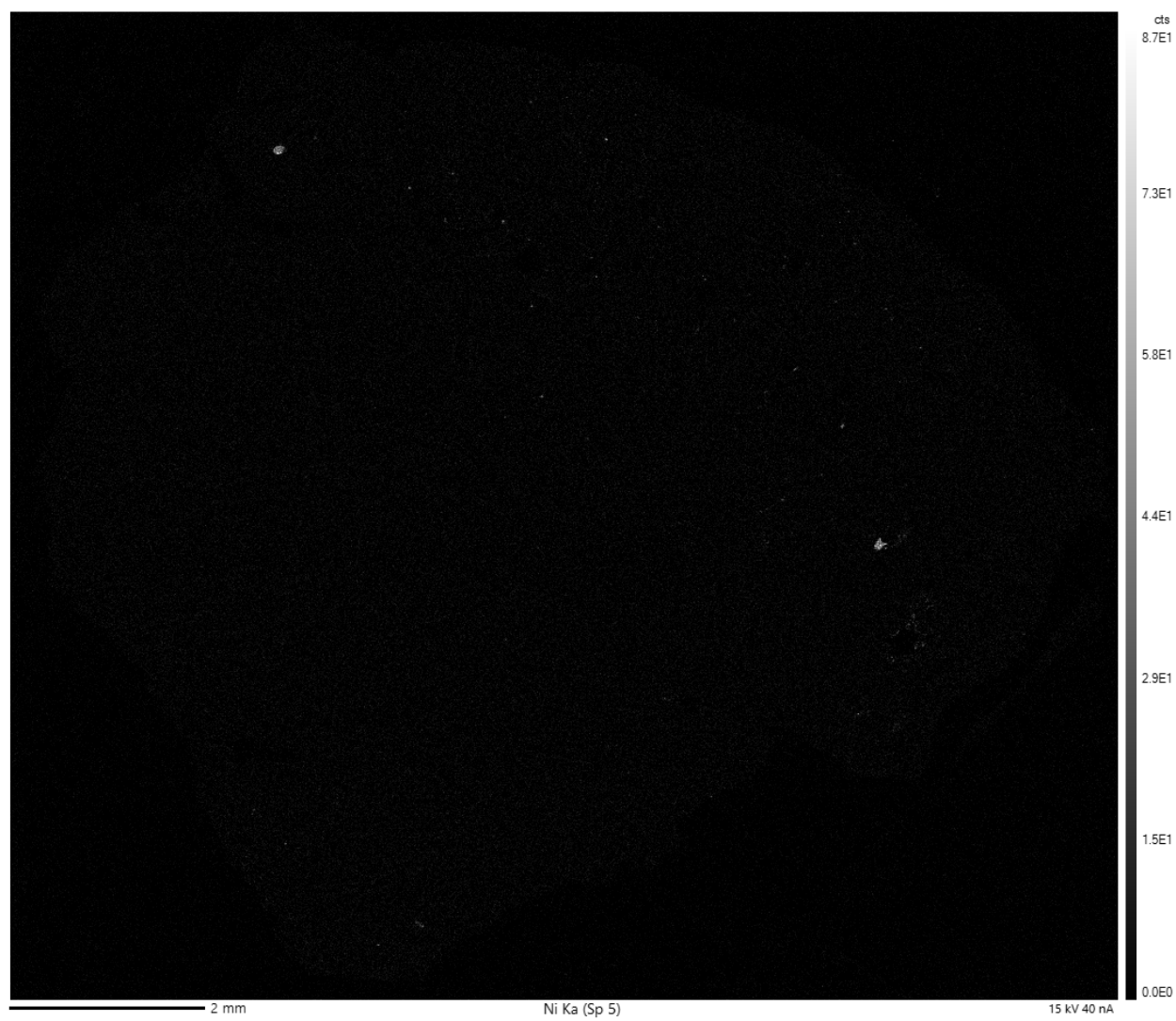


Figure S26. Ni X-ray map of 68815,148.

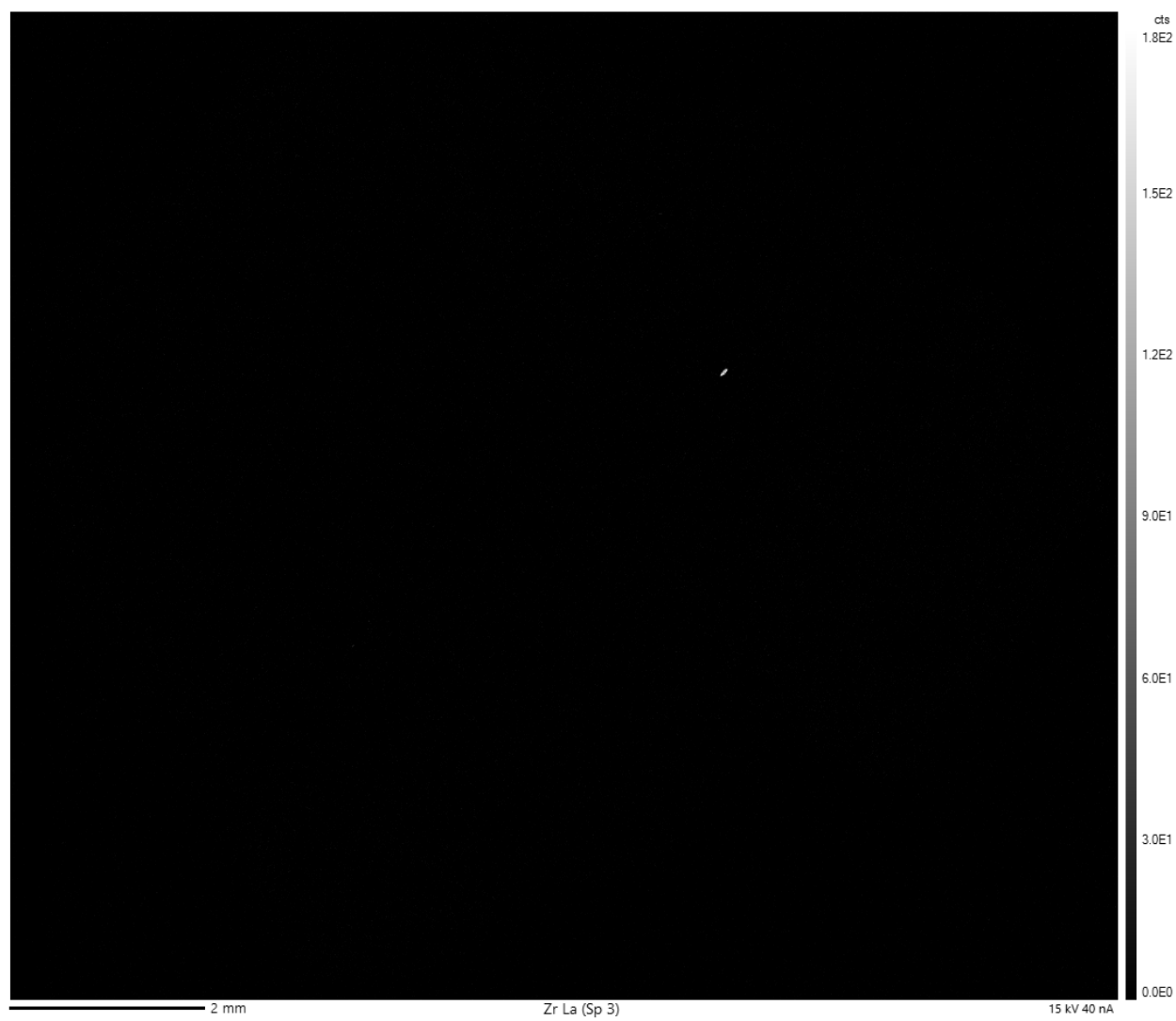


Figure S27. Zr X-ray map of 68815,148.

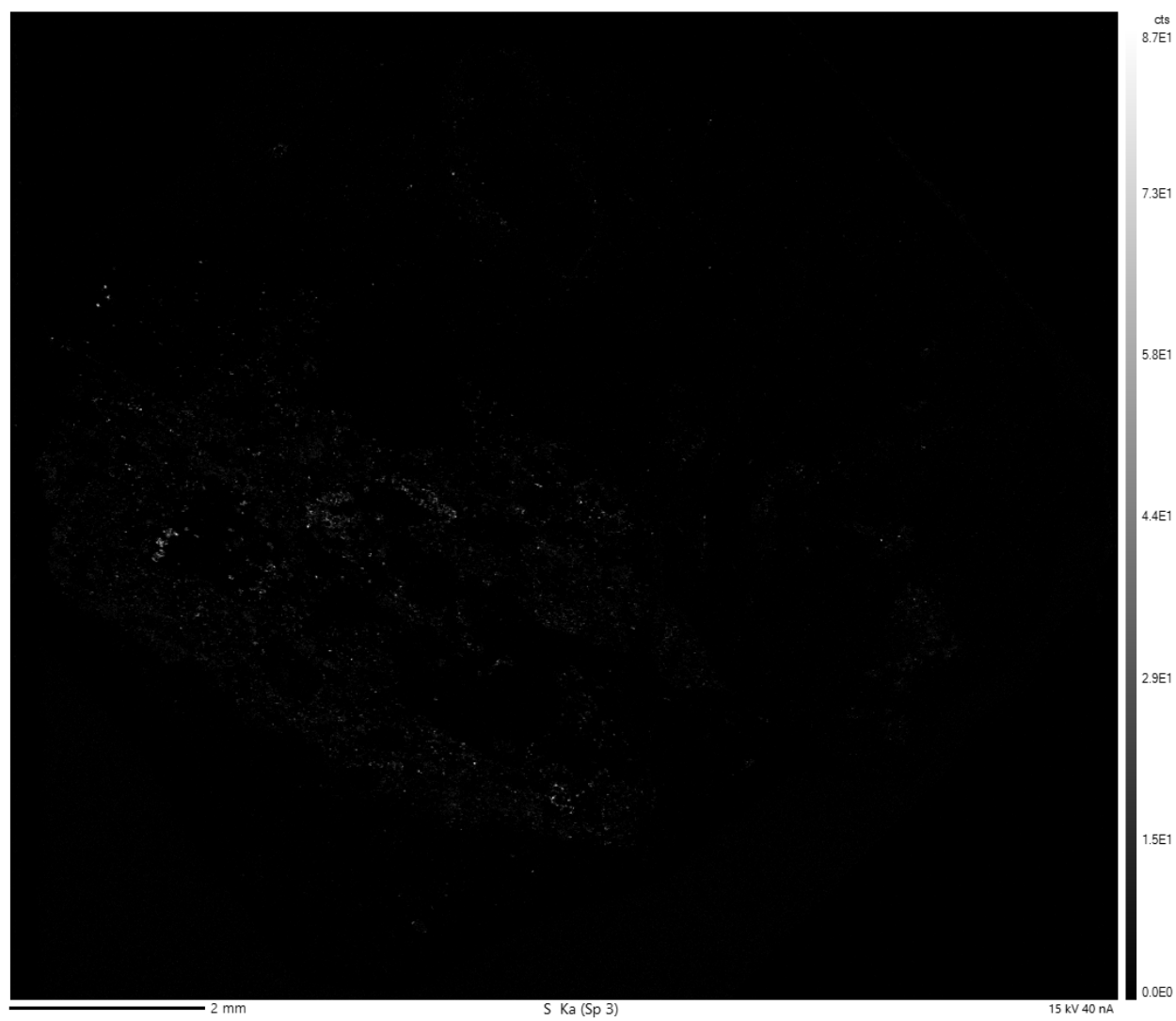


Figure S28. S X-ray map of 68815,148.

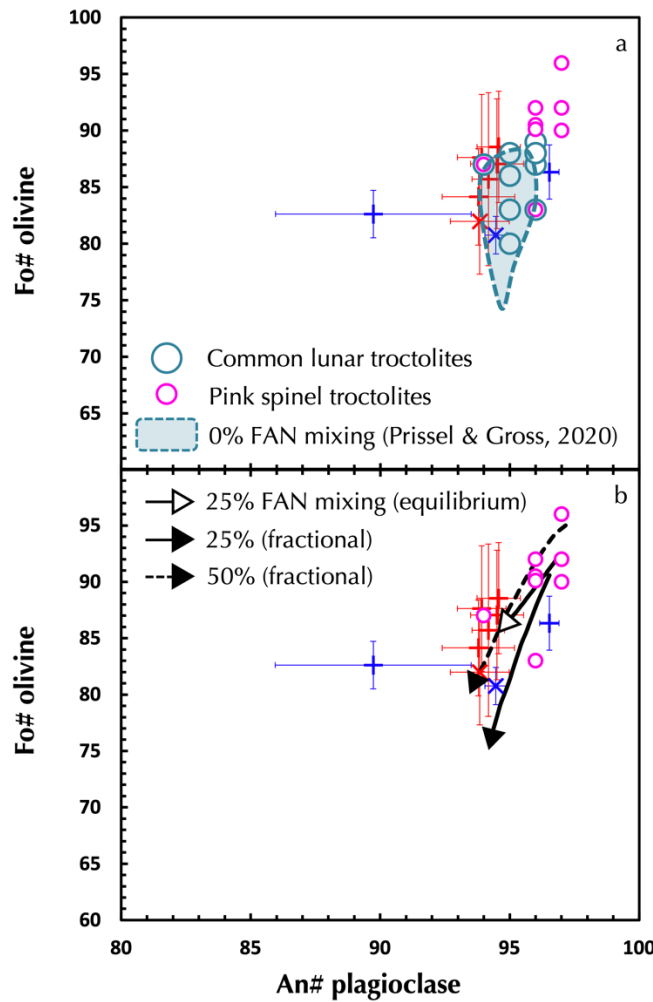


Figure S29. Forsterite content vs. anorthite content of co-existing plagioclase. (a) Symbols follow from previous figures with error bars encompassing the total observed variance relative to the median. Legend provided for common lunar troctolites ( $\pm$  chromite, no spinel) and pink spinel troctolites (*Shearer et al.*, 2015; and excluding only those data with undetermined values and evidence for metamorphism, i.e., cordierite). The combined equilibrium and fractional crystallization products from the mantle-derived Mg-suite primary melt from *Prissel and Gross* (2020) is shown and represents troctolitic assemblages produced with no crustal assimilation (i.e., common lunar troctolites). (b) Same as above, but now showing crystallization trends of the Mg-suite primary melt following 25 and 50% assimilation of lunar crust (equilibrium vs. fractional crystallization defined in the legend). Assimilation of crust by Mg-suite primary melts can explain the production of both pink spinel-bearing troctolites in addition to generally higher forsterite contents observed in the spinel troctolites. Additionally, the contaminated crystallization sequences provide linkages between the proposed basaltic clasts studied here to their plutonic equivalents in the Mg-suite. Further discussion provided in the main text.

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